

## *Quark-gluon plasma at RHIC and LHC*

## Electromagnetic and Rotational Fields at RHIC and LHC

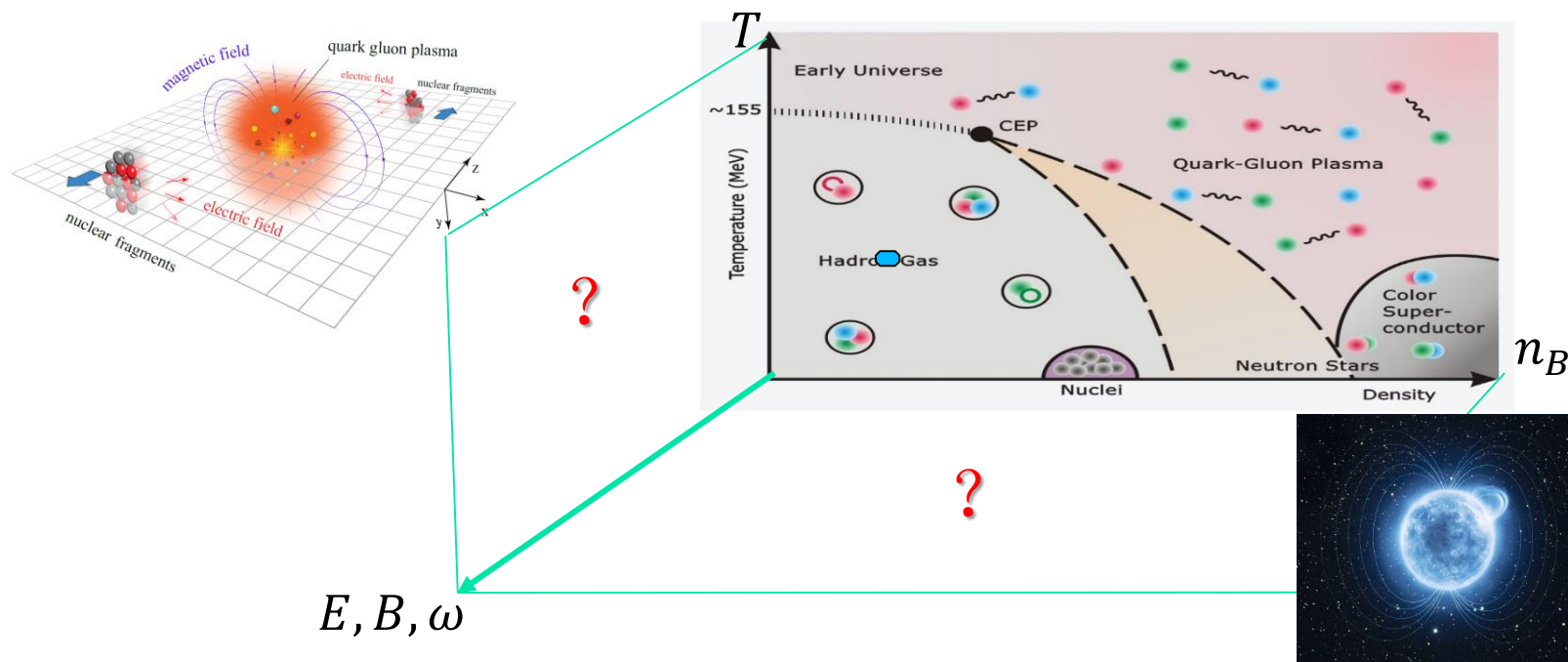
由于电磁相互作用远比强相互作用弱，一般研究QCD物质时不考虑电磁相互作用。

非对心核碰撞使QGP处于自然界最强的电磁场和涡旋场之中：

$$|eB| \sim 5m_\pi^2 \text{ at RHIC and } 70m_\pi^2 \text{ at LHC !}$$

$$\omega \sim 10^{21}/s \text{ at RHIC !}$$

其强度已经与强相互作用可比拟。



### 3 Questions

#### 1) Strength:

*E & B breaks down the translation invariance, but T will restore the invariance. Only charged quarks join electromagnetic interaction, but all the partons join thermal motion*

→ *Is the E and B field strong enough in comparison with the fireball T?*

$$|eB| \sim 10m_\pi^2, \quad T \sim 300 - 500 \text{ MeV}$$

#### 2) Time:

*The lifetime of external E and B field is very short, we need induced E and B field in QGP*

→ *Can the current be completely induced?*

$$\tau_{ind}, \quad \tau_B \sim 0.1 \text{ fm}$$

#### 3) Clear signals:

*If the answers to the above 2 questions are negative, is it possible to have sensitive electromagnetic signals of QCD matter?*

# Feynman Rules in Magnetic Field

Kostenko and Thompson, *Astrophys J.* 869, 44(2018), 875, 23(2019).

**External lines:**

$$[i\gamma^\mu(\partial_\mu + iqA_\mu) - m]\psi = 0$$

$$\psi_{\mp}^\sigma(\mathbf{x}, p) = \begin{cases} e^{-ip \cdot \mathbf{x}} u_\sigma(\mathbf{x}, p) \\ e^{ip \cdot \mathbf{x}} v_\sigma(\mathbf{x}, p) \end{cases}$$

$$u_-(\mathbf{x}, p) = \frac{1}{f_n} \begin{bmatrix} -ip_z p_n \phi_{n-1} \\ (\epsilon + \epsilon_n)(\epsilon_n + m)\phi_n \\ -ip_n(\epsilon + \epsilon_n)\phi_{n-1} \\ -p_z(\epsilon_n + m)\phi_n \end{bmatrix}, \quad v_+(\mathbf{x}, p) = \frac{1}{f_n} \begin{bmatrix} -p_n(\epsilon + \epsilon_n)\phi_{n-1} \\ -ip_z(\epsilon_n + m)\phi_n \\ -p_z p_n \phi_{n-1} \\ i(\epsilon + \epsilon_n)(\epsilon_n + m)\phi_n \end{bmatrix},$$

$$u_+(\mathbf{x}, p) = \frac{1}{f_n} \begin{bmatrix} (\epsilon + \epsilon_n)(\epsilon_n + m)\phi_{n-1} \\ -ip_z p_n \phi_n \\ p_z(\epsilon_n + m)\phi_{n-1} \\ ip_n(\epsilon + \epsilon_n)\phi_n \end{bmatrix}, \quad v_-(\mathbf{x}, p) = \frac{1}{f_n} \begin{bmatrix} -ip_z(\epsilon_n + m)\phi_{n-1} \\ -p_n(\epsilon + \epsilon_n)\phi_n \\ -i(\epsilon + \epsilon_n)(\epsilon_n + m)\phi_{n-1} \\ p_z p_n \phi_n \end{bmatrix}$$

**Quark propagator:**

$$G(x' - x) = -i \left( \frac{L}{2\pi\lambda} \right)^2 \int dp_z da \sum_{\sigma, n} \left[ \theta(t' - t) u_\sigma(\mathbf{x}', p) \bar{u}_\sigma(\mathbf{x}, p) e^{-ip \cdot (x' - x)} - \theta(t - t') v_\sigma(\mathbf{x}', p) \bar{v}_\sigma(\mathbf{x}, p) e^{ip \cdot (x' - x)} \right]$$

$$G(p) = - \int_0^\infty \frac{dv}{|qB|} \left\{ [m + (\gamma \cdot p)_\parallel] [1 - i \operatorname{sgn}(q) \gamma_1 \gamma_2 \tanh(v)] - \frac{(\gamma \cdot p)_\perp}{\cosh^2(v)} \right\} e^{-\frac{v}{|qB|} \left[ m^2 - p_\parallel^2 + \frac{\tanh(v)}{v} p_\perp^2 \right]}$$

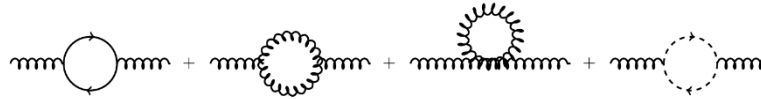
*Schwinger propagator, 1951*

● *no more translation invariance.*

● *the two Schwinger phases for  $q$  and  $\bar{q}$  are cancelled to each other in loop calculation.*

## Gluon Propagator in QCD Matter

*Gluon self-energy in magnetic field:*



*for quark loop*

$$\Pi_{\mu\mu}^{\parallel}(T, B) = g^2 T |qB| \sum_{np_z n_1} \frac{(2 - \delta_{n_1 0}) (\delta_{\mu\mu}^{\parallel} + g_{\mu\mu}^{\parallel}) (-\omega_n^2 + p_z^2)}{(m^2 + \omega_n^2 + p_z^2 + 2n_1 |qB|)^2}$$

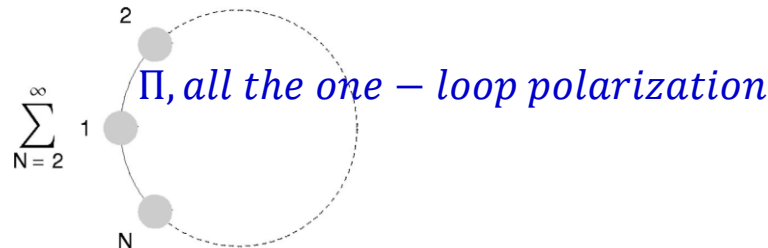
$$\Pi_{\mu\mu}^{\perp}(T, B) = 0$$

Matsubara frequencies  $p_0 = i\omega_n = i(2n + 1)\pi T$   
 quark longitudinal momentum  $p_z$   
 transverse Landau energy  $\varepsilon_k = 2n_1 |qB|$

*for gluon and ghost loops*

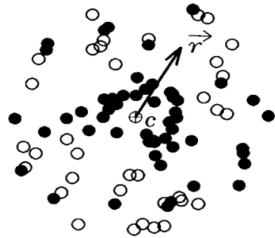
$$\bar{\Pi}_{\mu\nu}(T, B) = \bar{\Pi}_{\mu\nu}(T, 0)$$

*To include non-perturbative effect, we take the summation of ring diagrams  
 → gluon propagator and thermodynamic potential:*



## Color Screening in QCD Matter

Debye screening of a pair of charged particles  $q\bar{q}$ :



$$\frac{1}{r} \rightarrow \frac{1}{r} e^{-m_D r} = \frac{1}{r} e^{-r/r_D}$$

screening mass  $m_D$   
screening length  $r_D$

Pole of the propagator at gluon momentum ( $k_0^2 = 0$ ,  $\vec{k}^2 = -m_D^2$ )  $\rightarrow$  screening mass:

$$m_D^2(T, B) = m_Q^2(T, B) + m_G^2(T),$$

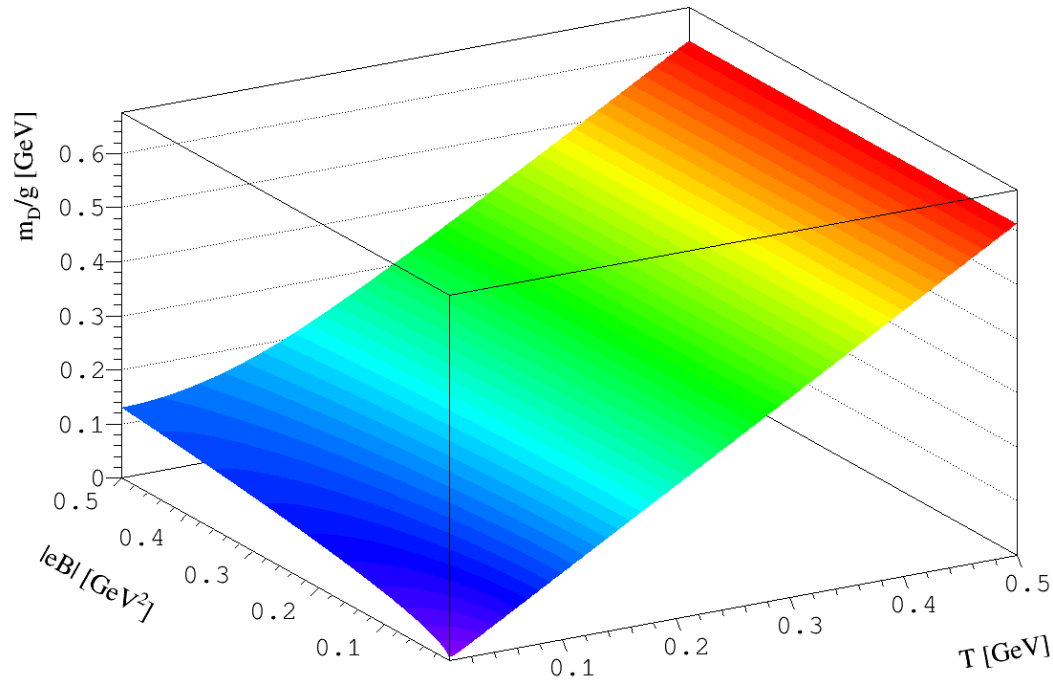
$$m_Q^2(T, B) = -\Pi_{00}^{\parallel}(T, B),$$

$$m_G^2(T) = -\bar{\Pi}_{00}^{\parallel}(T).$$

$$m_Q^2(T, B) = -g^2 T |qB| \sum_{np_z n_1} \left[ (2 - \delta_{n_1, 0}) \frac{m^2 - \omega_n^2 + p_z^2 + 2n_1 |qB|}{(m^2 + \omega_n^2 + p_z^2 + 2n_1 |qB|)^2} \right]$$

$$m_G^2(T) = \frac{N_c}{3} g^2 T^2$$

$$\underline{m_D(T, B)}$$



Huang, Zhao, Zhuang, PRD107, 114035(2023)

- $m_D(T = 0, B) = 0.13 \text{ GeV}$  at  $eB = 25m_\pi^2$ .
- the  $B$  effect is gradually washed out by thermal motion!

## Two time scales

- ♦ The induced current needs time  $\tau_{rel}$  to relax from zero to the Ohm's current

$$\vec{j}_{Ohm} = \sigma_{el}(\vec{E} + \vec{v} \times \vec{B})$$

- ♦ The lifetime of the external B-field:  $\tau_B$

*A complete electromagnetic response of QGP requires*

$$\tau_{rel} < \tau_B$$

*For normal materials like conductor and EM plasma,*

$$\tau_{rel} \ll \tau_B$$

*the relaxation time can be neglected.*

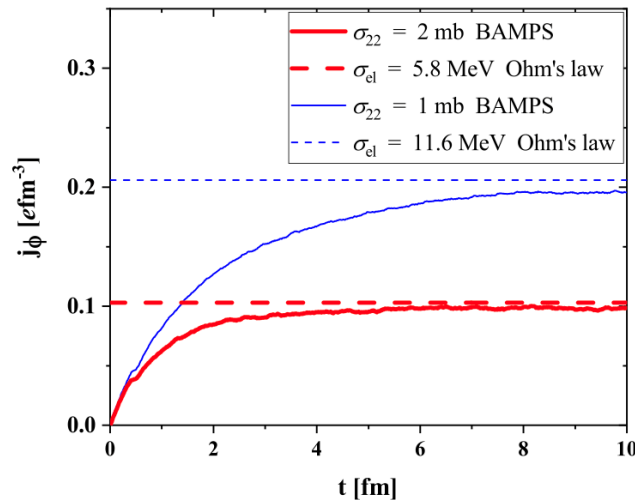
*How is about the relaxation in HIC where  $\tau_B$  and  $\tau_{QGP}$  are very short?*

## Broken Ohm's Law in HIC

The evolution of the fireball is described by a Boltzmann Approach of MultiParton Scatterings (BAMPS):

$$(\partial/\partial t + \vec{p}/E \cdot \vec{\nabla} + \vec{F} \cdot \vec{\nabla}_p)f = C_{22} + C_{23},$$

$$\text{Lorentz force } \vec{F} = q(\vec{p}/E \times \vec{B} + \vec{E})$$



Wang, Zhao, Greiner, Xu and Zhuang,  
PRC105, L041901(2022), Letter, Featured in Physics

$\vec{j} < \vec{j}_{Ohm}$ , Ohm's law is broken!

The electromagnetic response of the hot QCD matter to the fast decay of the external electromagnetic field is incomplete, which strongly suppresses the induced magnetic field.

## Color Screening in Dense QCD Matter

In the frame of ring diagram summation

$$m_D^2(\mu_f, B) = \frac{g^2}{(2\pi)^2} \sum_{n=0} (2 - \delta_{n0}) \frac{\mu_f |qB|}{\sqrt{\mu_f^2 - 2n|qB|}}$$

For chiral quarks, the distribution

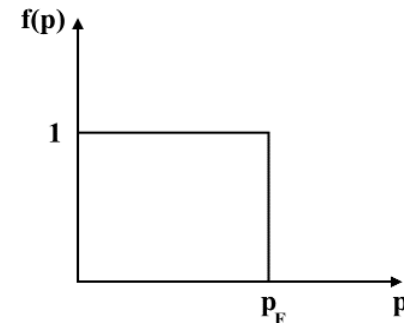
$$f(p) = \theta(\mu_f - p)$$

the Fermi surface is determined by

$$p_z^2 + 2n|qB| = \mu_f^2$$

when  $\mu_f$  and Landau levels are equal

$$\mu_f^2 = 2n|qB|$$

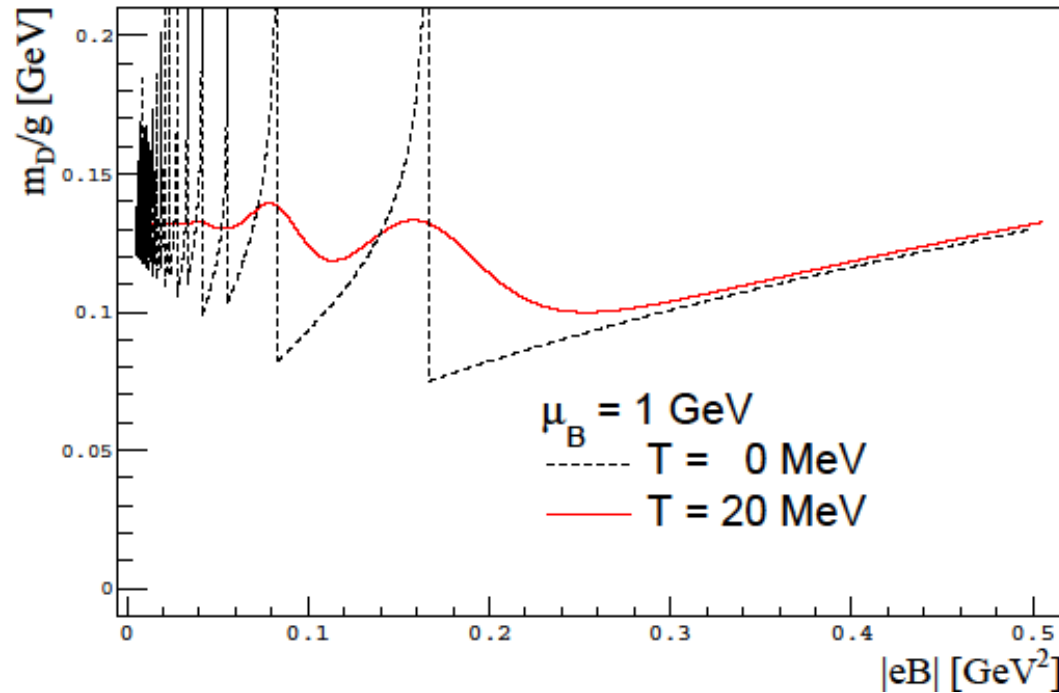


the infrared divergence at the Fermi surface induces a complete screening  $m_D \rightarrow \infty$ , called resonant screening.

The case here is similar to the resonant transmission in Quantum Mechanics.

## Resonant Screening

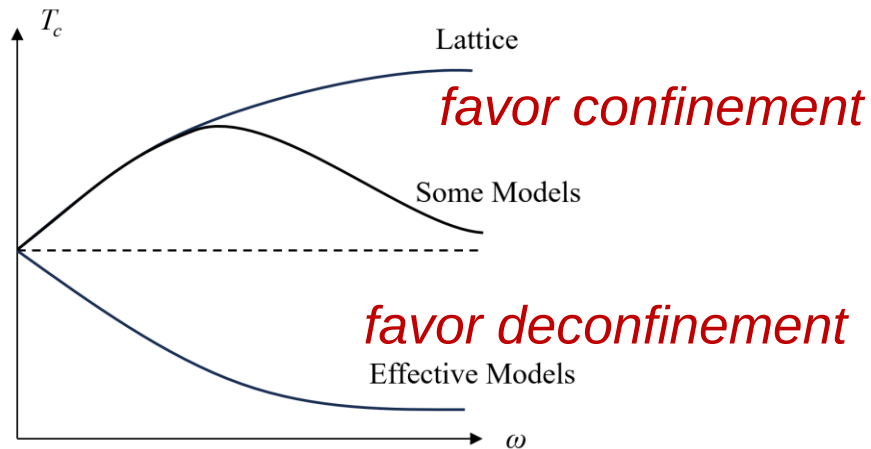
Huang, Zhao and Zhuang, PRD108, L091503(2023), Letter



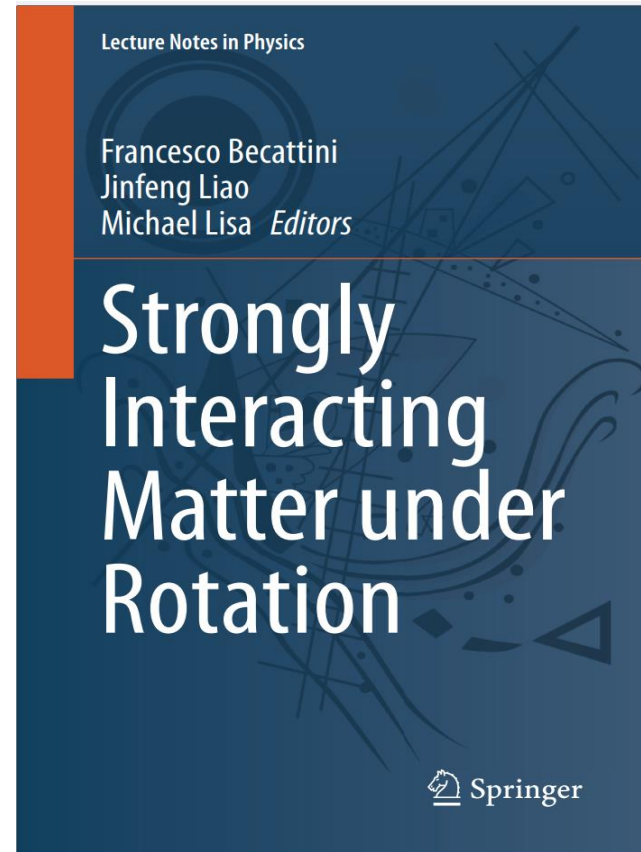
The color interaction between a pair of quarks is completely screened in dense and magnetized QCD matter, when chemical potential matches the Landau levels  $\mu_f^2 - 2n|qB| = 0$ .

## Rotation Puzzle

*Does a rotation favor QCD phase transition or not ?*



*We need serious calculations in QCD.*



## Gluon Propagator in Curved Space ( $\omega \neq 0$ )

*A rotating system is equivalent to a curved space,  
it can be described by a transformation between the flat and curved space.*

Flat space:  $\bar{x}_\mu$

Curved space:  $x_\mu$

$$\text{Transformation: } \begin{cases} t = \bar{t} \\ x = \bar{x} \cos(\omega \bar{t}) - \bar{y} \sin(\omega \bar{t}) \\ y = \bar{x} \sin(\omega \bar{t}) + \bar{y} \cos(\omega \bar{t}) \\ z = \bar{z} \end{cases}$$

Gluon self-energy:

$$\Pi_{\mu\nu}^{ab}(q, q') = \int \frac{dQ_0}{2\pi} \sum_{\vec{Q}} h_\mu^A(Q|q) h_\nu^B(-Q|q') \bar{\Pi}_{AB}^{ab}(Q)$$

*curved space*

*flat space*

$$h_\mu^A(Q|q) = \int d^4x \sqrt{-\det(g_{\sigma\rho}(x))} \frac{\partial \bar{x}^A}{\partial x^\mu} e^{i(qx - Q\bar{x})}$$

## Causality and Boundary Condition

**Causality:**

$$\omega R < 1$$

*For a rotating quark-gluon plasma created in RHIC,*

$$R \sim 10 \text{ fm}$$

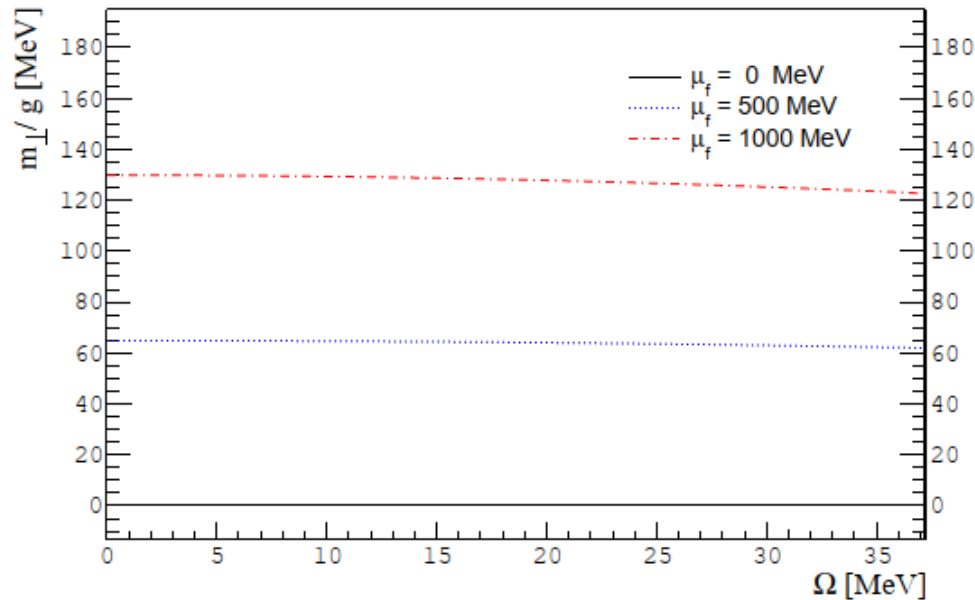
*$\omega \lesssim 30 \text{ MeV}$ , smaller in comparison with normal  $\mu_f$  and  $T$*

→ **Boundary condition:**

$$\begin{aligned}\vec{Q} &= \frac{2\pi}{L} \vec{N}, & \vec{N} &= (N_1, N_2, N_3) \\ \vec{q} &= \frac{2\pi}{L} \vec{n}, & \vec{n} &= (n_1, n_2, n_3)\end{aligned}$$

## Gluon Mass at Finite $\mu_f$ and $\omega$

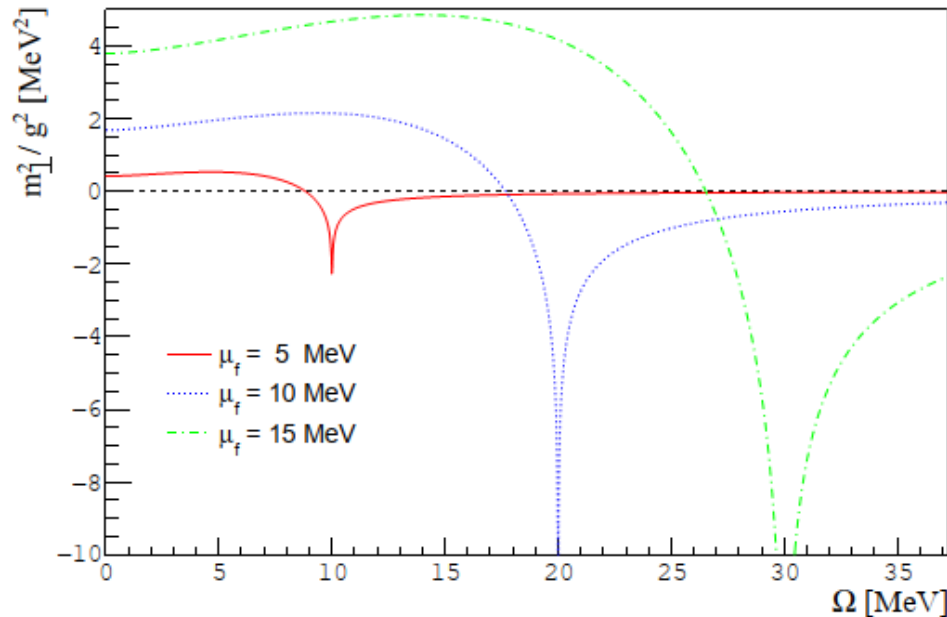
Huang, Chen, Jiang, Zhao and Zhuang, arXiv 2408\*\*\*



- 1)  $m$  decreases with  $\omega$ , it means an **antiscreening effect**.  
The density induced screening effect is partly cancelled by the rotation.  
**Rotation favors confinement.**
- 2)  $\omega$  ( $\lesssim 30$  MeV) is much smaller than  $\mu_f$  ( $\sim 500$  MeV), the rotation effect is weak, the maximum cancellation is  $\sim 5\%$ .
- 3) **There is no rotation effect in vacuum** ( $\mu_f = 0$ ).

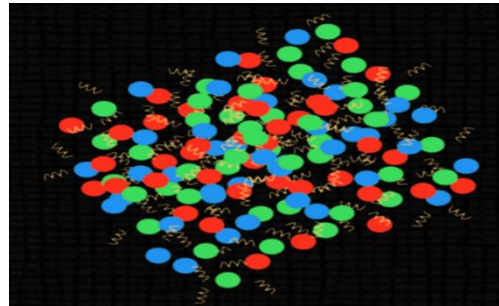
## Rotation Effect at Small $\mu_f$

Huang, Chen, Jiang, Zhao and Zhuang, arXiv 2408\*\*\*



- 1) At small  $\mu_f$ , the rotation effect becomes dominant!
- 2) The density effect is completely cancelled at a critical rotation  $\omega_c = 2\mu_f$ , and then the imaginary mass means that **gluons can no longer be considered as physical particles (confinement)**, as suggested by Gribov.
- 3) At finite  $T$ , the divergence is expected to become an oscillation.

## Classical Boltzmann equation



$$(p^\mu \partial_\mu^x + F^\mu \partial_\mu^p) f(x, p) = C$$

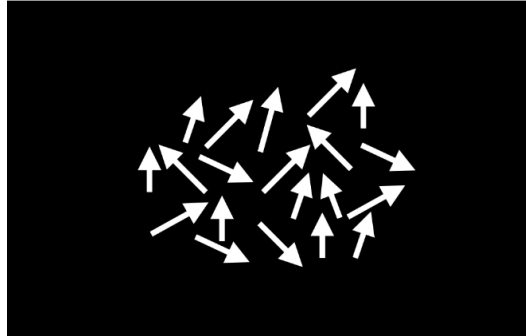
1) A single distribution (*particle number*)  $f$

2) *Quasi-particle* (on-shell) approximation:

$$(p^2 - m^2) f(x, p) = 0 \rightarrow f(x, \vec{p} | p_0 = \pm \sqrt{\vec{p}^2 + m^2})$$

## Spin

Many *quantum anomalies* in science are induced by spin, for instance CME and CVE.



### *Schrodinger equation*

*Wave function:*

*scalar*  $\psi(x)$

*Probability in coordinate space:*

$$\psi(x)\psi^*(x)$$

*Probability in phase space (Wigner transformation):*

$$f(x, p) = \int d^4 y e^{i p y} \psi(x + y/2) \psi^*(x - y/2)$$

### *Dirac equation*

$$\text{spinor } \psi(x) = \begin{pmatrix} \psi_1(x) \\ \psi_2(x) \\ \psi_3(x) \\ \psi_4(x) \end{pmatrix}$$

$$\psi(x) \bar{\psi}(x), \quad \text{a } 4 \times 4 \text{ matrix}$$

$$W(x, p) = \int d^4 y e^{i p y} \psi(x + y/2) \bar{\psi}(x - y/2)$$

*Wigner function with 16 distributions*

## Beyond quasi-particle

Particles in medium are in general not quasi-particles (*mean field approximation*), especially in high energy physics.

1) On-shell  $\rightarrow$  *Off-shell* constraint:

$$(p^2 - m^2)f(x, p) = 0 \rightarrow [(p^2 - m^2) + \hbar \mathcal{A}(p)]W(x, p) = 0$$

2) *Physics distributions*  $W(x, \vec{p})$  are defined in **7D** phase space  
 $\rightarrow$  *Equal-time formalism*

$$W(x, p) \rightarrow W(x, \vec{p}) = \int dp_0 W(x, p)$$

## Covariant and Equal-time Wigner functions

- *Covariant Wigner operator* for fermions interacting with a gauge field:

$$\widehat{W}(x, p) = \int d^4 y e^{i p y} \widehat{\psi}(x + \frac{y}{2}) e^{i q \int_{-1/2}^{1/2} ds \widehat{A}(x + s y) y} \widehat{\psi}(x - \frac{y}{2})$$

gauge link  $e^{i q \int_{-1/2}^{1/2} ds \widehat{A}(x + s y) y}$  to guarantee gauge invariance

*Ensemble average* → *Wigner function*

$$W(x, p) = \langle \widehat{W}(x, p) \rangle \quad (W = \widehat{W} \text{ in quantum mechanics})$$

- Dyson-Schwinger equation for quantum fields  $\widehat{\psi}$  or Dirac equation for wave function  $\psi$   
→ *kinetic equations for  $W(x, p)$*

*QED: D.Vasak, M.Gyulassy and H.-Th.Elze, Ann. Phys. 173, 462(1987)*

*QCD: H.-Th.Elze and U.Heinz, Phys. Rep. 183, 81(1989)*

- *Problem:*

Initial  $W(x, p)$  is related to the fields  $\widehat{\psi}(x)$  and  $\widehat{A}(x)$  at all times (due to  $\int_{-\infty}^{\infty} dy_0$ )

→ *Causality problem!*

*Physics distributions are defined through Equal-time Wigner function*

$$W_0(x, \vec{p}) = \int d^3 \vec{y} e^{-i \vec{p} \cdot \vec{y}} \left\langle \widehat{\psi}(x + \vec{y}/2) e^{-i q \int_{-1/2}^{1/2} ds \widehat{\vec{A}}(x + s \vec{y}) \cdot \vec{y}} \widehat{\psi}^+(x - \vec{y}/2) \right\rangle$$

## Dirac-Heisenberg-Wigner equation

Bialynicki-Birula, Gornicki and Rafelski, PRD44, 1825(1991)

### ● Dirac equation in coordinate space:

$$(i\gamma^\mu \mathcal{D}_\mu - m)\psi(x) = 0$$

→ DHW equation in phase space:

$$D_t W_0 = -\frac{1}{2} \vec{D} \cdot \{\rho_1 \vec{\sigma}, W_0\} - \frac{i}{\hbar} [\rho_1 \vec{\sigma} \cdot \vec{P} + \rho_3 m, W_0]$$

$$D_t = \partial_t + q \int_{-1/2}^{1/2} ds \vec{E}(\vec{x} + is\hbar \vec{V}_p) \cdot \vec{V}_p,$$

$$\vec{D} = \vec{V} + q \int_{-1/2}^{1/2} ds \vec{B}(\vec{x} + is\hbar \vec{V}_p) \times \vec{V}_p$$

$$\vec{P} = \vec{p} - iq\hbar \int_{-1/2}^{1/2} ds s \vec{B}(\vec{x} + is\hbar \vec{V}_p) \times \vec{V}_p$$

### ● However,

$$W_0(x, \vec{p}) = \int dp_0 W(x, p) \gamma_0$$

is not equivalent to  $W(x, p)$ . We should consider all the moments

$$W_n(x, \vec{p}) = \int dp_0 p_0^n W(x, p) \gamma_0 \quad (n = 0, 1, 2, \dots)$$

Zhuang and Heinz, Ann.Phys.245, 311(1996)

Only for quasi-particles (on-shell,  $p^2 - m^2 = 0$ ,  $p_0 = \pm \sqrt{m^2 + \vec{p}^2}$ ),

$$W_n(x, \vec{p}) = E_p^n W_0(x, \vec{p}),$$

$W_0(x, \vec{p})$  can completely describe the system.

## Equal-time hierarchy I

Zhuang and Heinz, PRD57, 6525(1998)

### 1) Covariant kinetic equations

$$\begin{aligned} (\gamma^\mu K_\mu - m)W &= 0 \\ K_\mu &= \Pi_\mu + \frac{i\hbar}{2} D_\mu & \Pi_\mu &= p_\mu - iq\hbar \int_{-1/2}^{1/2} ds s F_{\mu\nu}(x - i\hbar s \partial_p) \partial_p^\nu \\ & & D_\mu &= \partial_\mu - q \int_{-\frac{1}{2}}^{\frac{1}{2}} ds F_{\mu\nu}(x - i\hbar s \partial_p) \partial_p^\nu \end{aligned}$$

### Constraint (with $p_\mu$ ) and transport (with $\partial_\mu$ ) equations

$$\begin{cases} [\gamma^\mu (K_\mu + K_\mu^\dagger) - 2m]W(x, p) = 0 \\ \gamma^\mu (K_\mu - K_\mu^\dagger)W(x, p) = 0 \end{cases}$$

### 2) Equal-time hierarchy

$$\int dp_0 [\text{covariant constraint and transport equations}] \rightarrow$$

$$\begin{cases} \text{Transport equations for } W_0(x, \vec{p}) & \rightarrow \text{DHW equation} \\ \text{Constraint equation for } W_1(x, \vec{p}) \end{cases}$$

$$\int dp_0 p_0 \cdot [\text{covariant constraint and transport equations}] \rightarrow$$

$$\begin{cases} \text{Transport equations for } W_1(x, \vec{p}) \\ \text{Constraint equation for } W_2(x, \vec{p}) \end{cases}$$

## Equal-time hierarchy II

*Zhuang and Heinz, PRD57, 6525(1998)*

### 3) Spin decomposition

$$W_0(x, \vec{p}) = \frac{1}{4} [f_0 + \gamma^5 f_1 - i\gamma^0 \gamma^5 f_2 + \gamma^0 f_3 + \gamma^5 \gamma^0 \vec{\gamma} \cdot \vec{g}_0 + \gamma^0 \vec{\gamma} \cdot \vec{g}_1 - i\vec{\gamma} \cdot \vec{g}_2 - \gamma^5 \vec{\gamma} \cdot \vec{g}_3]$$

*D.Vasak, M.Gyulassy and H.-Th.Elze, Ann. Phys. 173, 462(1987)*

### Conservation laws → Physics of the spin components

$f_0$  : number density,       $\vec{g}_0$  : spin density  
 $f_1$  : helicity density,       $f_2$  : topologic charge density,       $f_3$  : mass density  
 $\vec{g}_1$  : number current,       $\vec{g}_3$  : magnetic moment

### 4) Truncating the hierarchy

spin 1/2 particles:  $W_0$  and  $W_1$  form a closed subgroup

spin 0 particles:  $W_0, W_1$  and  $W_2$  form a closed subgroup

## Spin distribution at lowest level

Zhuang and Heinz, PRD53, 2096(1996)

$$W_0(x, \vec{p}) = W_0^{(0)}(x, \vec{p}) + \hbar W_0^{(1)}(x, \vec{p}) + \dots$$

- *Constraint equations → only 4 independent components:*  
number density  $f_0$  and spin density  $\vec{g}_0$

- *Boltzmann equation for number density  $f_0$ :*

$$\left( D_t + \frac{\vec{p}}{E_p} \cdot \vec{D} \right) f_0 = 0$$

$$D_t = \partial_t + q \vec{E} \cdot \vec{\nabla}_p, \quad \vec{D} = \vec{V} + q \vec{B} \times \vec{\nabla}_p$$

- *Bargmann-Michel-Telegdi equation for spin density  $\vec{g}_0$ :*

$$\left( D_t + \frac{\vec{p}}{E_p} \cdot \vec{D} \right) \vec{g}_0 = \frac{q}{E_p^2} [\vec{p} \times (\vec{E} \times \vec{g}_0) - E_p \vec{B} \times \vec{g}_0]$$

*the phase-space version of the Bargmann-Michel-Telegdi equation*

*V.Bargmann, L.Michel and V.Telegdi, PRL2, 435(1959)*

## Quark Matter in a Rotation Field

Chen and Zhuang, CPC (2022)

Dirac equation:

$$(i\gamma^\mu \partial_\mu - m + \gamma_0 \vec{\omega} \cdot \hat{j})\psi(x) = 0$$

Covariant Wigner function:

$$W(x, p) = \int \frac{d^4 y}{(2\pi)^4} e^{ip \cdot y} \left\langle \psi\left(x + \frac{y}{2}\right) \bar{\psi}\left(x - \frac{y}{2}\right) \right\rangle$$

Covariant kinetic equation:

$$\left( \gamma^\mu K_\mu + \frac{\hbar}{2} \gamma^5 \gamma^\mu \omega_\mu - m \right) W(x, p) = 0$$

$$K_\mu = \Pi_\mu + \frac{i\hbar}{2} D_\mu, \quad \omega_\mu = (0, \vec{\omega})$$

$$\Pi_\mu = (p_0 + \pi_0, \vec{p}), \quad \pi_0 = \vec{\omega} \cdot \left( \vec{l} + \frac{\hbar^2}{4} \vec{V} \times \vec{V}_p \right) + \mu_B$$

$$D_\mu = (d_t, \vec{V}), \quad d_t = \partial_t - \vec{\omega} \cdot (\vec{x} \times \vec{V} + \vec{p} \times \vec{V}_p)$$

## Classical Transport in a Rotation Field

At order  $\hbar^0$ , only  $f_0$  (number distribution) and  $\vec{g}_0$  (spin distribution) are independent components.

16 transport equations are reduced to

Chen and Zhuang, CPC (2022)

$$\left[ \partial_t + \left( \pm \frac{\mathbf{p}}{\epsilon_p} + \mathbf{x} \times \boldsymbol{\omega} \right) \cdot \nabla - (\boldsymbol{\omega} \times \mathbf{p}) \cdot \nabla_p \right] f_0^{(0)\pm} = 0,$$

$$\left[ \partial_t + \left( \pm \frac{\mathbf{p}}{\epsilon_p} + \mathbf{x} \times \boldsymbol{\omega} \right) \cdot \nabla - (\boldsymbol{\omega} \times \mathbf{p}) \cdot \nabla_p \right] \mathbf{g}_0^{(0)\pm} = -\boldsymbol{\omega} \times \mathbf{g}_0^{(0)\pm}.$$

Coriolis force     spin-rotation coupling

16 constraint equations are reduced to

on-shell energy  $E_p^\pm = \pm \epsilon_p - (\vec{\omega} \cdot \vec{l} + \mu_B), \quad \epsilon_p = \sqrt{m^2 + p^2}$

$$f_1^{(0)\pm} = \pm \frac{1}{\epsilon_p} \mathbf{p} \cdot \mathbf{g}_0^{(0)\pm},$$

$$f_2^{(0)\pm} = 0,$$

$$f_3^{(0)\pm} = \pm \frac{m}{\epsilon_p} f_0^{(0)\pm},$$

$$\mathbf{g}_1^{(0)\pm} = \pm \frac{\mathbf{p}}{\epsilon_p} f_0^{(0)\pm},$$

$$\mathbf{g}_2^{(0)\pm} = \frac{1}{m} \mathbf{p} \times \mathbf{g}_0^{(0)\pm},$$

$$\mathbf{g}_3^{(0)\pm} = \pm \frac{1}{m\epsilon_p} \left[ \epsilon_p^2 \mathbf{g}_0^{(0)\pm} - \mathbf{p}(\mathbf{p} \cdot \mathbf{g}_0^{(0)\pm}) \right].$$

One can systematically calculate the higher order contributions.

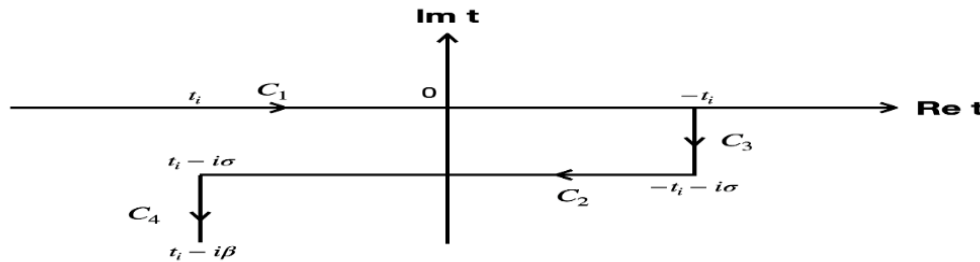
## Kadanoff-Baym equations

### Dyson-Schwinger equation

$$S(x, y) = S^0(x, y) + \int d^4z d^4w S^0(x, w) \Sigma(w, z) S(z, y)$$

Considering the time order in  $W(x, p) = \int d^4y e^{ipy} \psi\left(x + \frac{y}{2}\right) \bar{\psi}\left(x - \frac{y}{2}\right)$

→ Schwinger-Keldish time contour



$$S^<(x, p) = W(x, p), \quad S^>(x, p)$$

### Constraint and transport equations with collision terms

$$\begin{aligned} \{(\gamma^\mu p_\mu - m), S^<\} + \frac{i\hbar}{2} [\gamma^\mu, \nabla_\mu S^<] &= \frac{i\hbar}{2} ([\Sigma^<, S^>]_* - [\Sigma^>, S^<]_*) \\ [(\gamma^\mu p_\mu - m), S^<] + \frac{i\hbar}{2} \{\gamma^\mu, \nabla_\mu S^<\} &= \frac{i\hbar}{2} (\{\Sigma^<, S^>\}_* - \{\Sigma^>, S^<\}_*) \\ A * B &= AB + \frac{i\hbar}{2} [AB]_{P.B.} + \mathcal{O}(\hbar^2) \end{aligned}$$

including spin: Yang, Hattori, Hidaka, JHEP 2020 (2020) 070, arXiv:2002.02612

### ● Spin decomposition for Wigner function, self-energy and collision term

# General collision terms

Wang, Guo and Zhuang, EPJC, (2021)

## 0th order transport

描述相对论重离子碰撞的许多模型, 例如URQMD, AMPT, BAMPS.

$$p \cdot \nabla \mathcal{V}_\mu^{(0)} = m \widehat{\Sigma_S^{(0)} \mathcal{V}_\mu^{(0)}} + p^\nu \widehat{\Sigma_{V\nu}^{(0)} \mathcal{V}_\mu^{(0)}} + \frac{m}{2} \epsilon_{\alpha\beta\lambda\mu} \widehat{\Sigma_T^{(0)\alpha\beta} \mathcal{A}^{(0)\lambda}} - \frac{p_\nu}{m} \epsilon_{\alpha\mu\beta\lambda} p^\beta \widehat{\Sigma_T^{(0)\alpha\nu} \mathcal{A}^{(0)\lambda}} - p_\mu \widehat{\Sigma_A^{(0)\nu} \mathcal{A}_\nu^{(0)}},$$

$$p \cdot \nabla \mathcal{A}_\mu^{(0)} = m \widehat{\Sigma_S^{(0)} \mathcal{A}_\mu^{(0)}} + p^\nu \widehat{\Sigma_{V\nu}^{(0)} \mathcal{A}_\mu^{(0)}} + \frac{m}{2} \epsilon_{\alpha\beta\lambda\mu} \widehat{\Sigma_T^{(0)\alpha\beta} \mathcal{V}^{(0)\lambda}} + \widehat{\Sigma_{A\mu}^{(0)} p^\nu \mathcal{V}_\nu^{(0)}} - p_\mu \widehat{\Sigma_{A\nu}^{(0)} \mathcal{V}^{(0)\nu}},$$

$$\widehat{FG} = \bar{F}G - F\bar{G}$$

Local collision term  
Dynamical effect,  
e.g. diffusion effect

## 1st order transport

$$p \cdot \nabla \mathcal{V}_\mu^{(1)} = + m \widehat{\Sigma_S^{(0)} \mathcal{V}_\mu^{(1)}} + p^\nu \widehat{\Sigma_{V\nu}^{(0)} \mathcal{V}_\mu^{(1)}} - p_\mu \widehat{\Sigma_A^{(0)\nu} \mathcal{A}_\nu^{(1)}} - \frac{p_\nu}{m} \epsilon_{\rho\sigma\alpha\mu} p^\rho \widehat{\Sigma_T^{(0)\alpha\nu} \mathcal{A}^{(1)\sigma}} + \frac{m}{2} \epsilon_{\sigma\nu\lambda\mu} \widehat{\Sigma_T^{(0)\sigma\nu} \mathcal{A}^{(1)\lambda}}$$

$$+ m \widehat{\Sigma_S^{(1)} \mathcal{V}_\mu^{(0)}} + p^\nu \widehat{\Sigma_{V\nu}^{(1)} \mathcal{V}_\mu^{(0)}} - p_\mu \widehat{\Sigma_A^{(1)\nu} \mathcal{A}_\nu^{(0)}} - \frac{p_\nu}{m} \epsilon_{\alpha\mu\beta\lambda} p^\beta \widehat{\Sigma_T^{(1)\alpha\nu} \mathcal{A}^{(0)\lambda}} + \frac{m}{2} \epsilon_{\sigma\nu\lambda\mu} \widehat{\Sigma_T^{(1)\sigma\nu} \mathcal{A}^{(0)\lambda}}$$

$$+ \frac{1}{2m} p^\nu [\widehat{\Sigma_{T\mu\nu}^{(0)} (p^\alpha \mathcal{V}_\alpha^{(0)})}]_{\text{P.B.}} - \frac{m}{2} [\widehat{\Sigma_{T\mu\nu}^{(0)} \mathcal{V}^{(0)\nu}}]_{\text{P.B.}} + \frac{1}{2} \epsilon_{\mu\nu\alpha\beta} p^\nu [\widehat{\Sigma_A^{(0)\alpha} \mathcal{V}^{(0)\beta}}]_{\text{P.B.}}$$

$$- \frac{1}{2m} p_\nu \widehat{\Sigma_{T\alpha\mu}^{(0)} \nabla^\alpha \mathcal{V}^{(0)\nu}} + \frac{1}{2m} p_\nu \widehat{\Sigma_T^{\alpha\nu(0)} \nabla_{[\alpha} \mathcal{V}_{\mu]}^{(0)}} + \frac{1}{2} \epsilon_{\beta\nu\lambda\mu} \widehat{\Sigma_A^{(0)\beta} \nabla^\nu \mathcal{V}^{(0)\lambda}}$$

$$+ \frac{1}{2} \epsilon_{\mu\nu\alpha\beta} (\nabla^\alpha \widehat{\Sigma_V^{(0)\nu}}) \mathcal{A}^{(0)\beta} - \frac{1}{2m} p_\mu (\nabla^\nu \widehat{\Sigma_P^{(0)}}) \mathcal{A}_\nu^{(0)} - \frac{1}{2m} (p^\nu \nabla_\nu \widehat{\Sigma_P^{(0)}}) \mathcal{A}_\mu^{(0)} + \frac{p^\nu}{2m} \epsilon_{\mu\nu\alpha\beta} (\nabla^\alpha \widehat{\Sigma_S^{(0)}}) \mathcal{A}^{(0)\beta},$$

$$p \cdot \nabla \mathcal{A}_\mu^{(1)} = + m \widehat{\Sigma_S^{(0)} \mathcal{A}_\mu^{(1)}} + p^\nu \widehat{\Sigma_{V\nu}^{(0)} \mathcal{A}_\mu^{(1)}} + p^\nu \widehat{\Sigma_{A\mu}^{(0)} \mathcal{V}_\nu^{(1)}} + \frac{m}{2} \epsilon_{\alpha\beta\lambda\mu} \widehat{\Sigma_T^{(0)\alpha\beta} \mathcal{V}^{(1)\lambda}} - p_\mu \widehat{\Sigma_{A\nu}^{(0)} \mathcal{V}^{(1)\nu}}$$

$$+ m \widehat{\Sigma_S^{(1)} \mathcal{A}_\mu^{(0)}} + p^\nu \widehat{\Sigma_{V\nu}^{(1)} \mathcal{A}_\mu^{(0)}} + p^\nu \widehat{\Sigma_{A\mu}^{(1)} \mathcal{V}_\nu^{(0)}} + \frac{m}{2} \epsilon_{\alpha\beta\lambda\mu} \widehat{\Sigma_T^{(1)\alpha\beta} \mathcal{V}^{(0)\lambda}} - p_\mu \widehat{\Sigma_{A\nu}^{(1)} \mathcal{V}^{(0)\nu}}$$

$$- \frac{1}{2} \epsilon_{\mu\nu\rho\sigma} (\nabla^\sigma \widehat{\Sigma_V^{(0)\nu}}) \mathcal{V}^{(0)\rho} - \frac{m}{2} [\widehat{\Sigma_P^{(0)} \mathcal{V}_\mu^{(0)}}]_{\text{P.B.}} + \frac{1}{2m} p_\mu [\widehat{\Sigma_P^{(0)} (p^\nu \mathcal{V}_\nu^{(0)})}]_{\text{P.B.}}$$

$$+ \frac{1}{2} \epsilon_{\mu\sigma\nu\rho} \nabla^\sigma \widehat{\Sigma_A^{(0)\nu} \mathcal{A}^{(0)\rho}} + \frac{1}{2} \epsilon_{\nu\mu\alpha\beta} [\widehat{\Sigma_A^{(0)\nu} (p^\alpha \mathcal{A}^{(0)\beta})}]_{\text{P.B.}} - \frac{m}{2} [\widehat{\Sigma_{T\mu\nu}^{(0)} \mathcal{A}^{(0)\nu}}]_{\text{P.B.}} + \frac{p_\mu}{2m} [\widehat{\Sigma_{T\rho\nu}^{(0)} (p^\rho \mathcal{A}^{(0)\nu})}]_{\text{P.B.}}$$

$$- \frac{1}{2m} p_\sigma \nabla^\sigma (\widehat{\Sigma_{T\mu\nu}^{(0)} \mathcal{A}^{(0)\nu}}) + \frac{1}{2m} p^\nu \nabla^\sigma (\widehat{\Sigma_{T\mu\nu}^{(0)} \mathcal{A}_\sigma^{(0)}}) + \frac{1}{2m} p_\mu \nabla^\sigma (\widehat{\Sigma_{T\sigma\nu}^{(0)} \mathcal{A}^{(0)\nu}}) - \frac{1}{2m} p^\nu \nabla^\sigma (\widehat{\Sigma_{T\sigma\nu}^{(0)} \mathcal{A}_\mu^{(0)}}).$$

- Nonlocal collision term
- Related to spatial derivatives
- Correlated transport of V&A
- Polarization can be generated in a initially unpolarized system

the interaction needs to be specified to calculate the off-diagonal self-energy  $\Sigma^>$  &  $\Sigma^<$

## Near equilibrium state

Z.Wang and P.Zhuang, [arXiv:2105.00915 [hep-ph]].

- *Collision* is the driving force for the system to go from *non-equilibrium* to equilibrium.
- *Local equilibrium state* is determined by *detailed balance* principle (loss term and gain term cancel to each other).
- *Global equilibrium state* is determined by *detailed balance + Killing condition*

$$\partial_\mu \beta_\nu + \partial_\nu \beta_\mu = 0, \quad \beta_\mu = u_\mu/T$$

- *Near-equilibrium state* can be described by *relaxation time approximation (RTA)*.

## Relaxation Time Approximation (RTA)

Classical RTA: [J.Anderson and H.Witting, Physica 74, 466(1974)]

$$p^\mu \partial_\mu f = -p^\mu u_\mu \frac{\delta f}{\tau}, \quad \delta f = f - f_{eq}$$

Quantum RTA:

$$(\gamma^\mu p_\mu - m)W + \frac{i\hbar}{2} \gamma^\mu \partial_\mu W = -\frac{i\hbar}{2} \gamma^\mu u_\mu \frac{\delta W}{\tau}, \quad \delta W = W - W_{eq}$$

*A single time scale to control the relaxation process for all the components.*

Motivations:

- 1) A multi-component kinetic theory needs more time scales to control different degrees of freedom.*
- 2) Calculating the relaxation times in Kadanoff-Baym formalism.*

## RTA from Kadanoff-Baym equations

Z.Wang and P.Zhuang, [arXiv:2105.00915 [hep-ph]].

Calculating *damping and correlation times* for different degrees of freedom with Kadanoff-Baym equations.

### *Method:*

1) Quantum kinetic equations in near equilibrium state:

$$W = W_{eq} + \delta W$$

2) Expanding collision terms around the equilibrium state:

$$C(\Sigma, W) = C'(\Sigma_{eq}, W_{eq})\delta W$$

Note:  $C(\Sigma_{eq}, W_{eq}) = 0$  in equilibrium state

### *Kadanoff-Baym RTA:*

Chiral fermions:  $f_{\pm} = \frac{1}{2}(f_V \pm f_A)$

$$\begin{cases} p^{\mu} \partial_{\mu} f_{+} = -\frac{\delta f_{+}}{\tau_{+}} - \frac{\delta f_{+} - \delta f_{-}}{\tau_{+-}} \\ p^{\mu} \partial_{\mu} f_{-} = -\frac{\delta f_{-}}{\tau_{-}} - \frac{\delta f_{-} - \delta f_{+}}{\tau_{+-}} \end{cases}$$

$\tau_{\pm}$ : damping time for chiral fermions

$\tau_{+-}$ : correlation time between two kinds of fermions